Blending Data And Models: Kalman Based Approaches

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Overview

Historical Context: Celestial Mechanics

Kalman Methodology & Applications

Weather Forecasting

Ensemble Kalman Inversion

Gradient Flow

Closing



Historical Context: Celestial Mechanics

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Brahe



Purely observational data – initially by eye. Big data c. 1600s.

Kepler



Mathematical formulae which interpolated Brahe's data. Data-driven model: Kepler's Law.

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Newton



Kepler's Law rationalized through Newtonian mechanics. Led to theory of conservation laws: extrapolation.

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Einstein



Discrepancy between data and predictions of Newtonian mechanics. Mercury perehilion; resolved by special and then general relativity. The scientific method.

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Kalman Methodology & Applications

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Kalman Filter (Navigation)

State Space Model

Dynamics Model: $v_{n+1} = Mv_n + \xi_n$, $n \in \mathbb{Z}^+$ Data Model: $y_{n+1} = Hv_{n+1} + \eta_{n+1}$, $n \in \mathbb{Z}^+$ Probabilistic Structure: $v_0 \sim N(m_0, C_0)$, $\xi_n \sim N(0, \Sigma)$, $\eta_n \sim N(0, \Gamma)$ Probabilistic Structure: $v_0 \perp \{\xi_n\} \perp \{\eta_n\}$ independent



- Rudolph Kalman.
- J. Basic Engineering 82(1960); [19].
- ▶ \approx 35,000 citations (Google Scholar 9/20).

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- Apollo 11.
- $\triangleright \quad Y_n = \{y_\ell\}_{\ell=1}^n.$
- $\triangleright v_n | Y_n \sim N(m_n, C_n).$
- $\blacktriangleright (m_n, C_n) \mapsto (m_{n+1}, C_{n+1}).$

Kalman Filter

Sequential Optimization Perspective

Predict:
$$\widehat{m}_{n+1} = Mm_n$$
, $n \in \mathbb{Z}^+$
Model/Data Compromise: $J_n(m) = \frac{1}{2}|m - \widehat{m}_{n+1}|^2_{\widehat{C}_{n+1}} + \frac{1}{2}|y_{n+1} - Hm|^2_{\Gamma}$
Optimize: $m_{n+1} = \operatorname{argmin}_m J_n(m)$.

•
$$|\cdot|_A = |A^{-\frac{1}{2}} \cdot |$$
 for $A > 0$.

- *d* the state space dimension $(m_n, v_n \in \mathbb{R}^d)$.
- Updating \widehat{C}_{n+1} is expensive: $\mathcal{O}(d^2)$ storage.

3DVAR Filter (Weather Forecasting)

State Space Model

Dynamics Model: $v_{n+1} = \Psi(v_n) + \xi_n$, $n \in \mathbb{Z}^+$ Data Model: $y_{n+1} = Hv_{n+1} + \eta_{n+1}$, $n \in \mathbb{Z}^+$ Probabilistic Structure: $v_0 \sim N(m_0, C_0)$, $\xi_n \sim N(0, \Sigma)$, $\eta_n \sim N(0, \Gamma)$ Probabilistic Structure: $v_0 \perp \{\xi_n\} \perp \{\eta_n\}$ independent



- Andrew Lorenc.
- Introduced in UK Met Office.

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- See [25]
- $\blacktriangleright \{v_n\} \mapsto \{v_{n+1}\}.$

3DVAR

Sequential Optimization Perspective

$$\begin{array}{ll} {\sf Predict:} & \widehat{v}_{n+1} = \Psi(v_n), & n \in \mathbb{Z}^+\\ {\sf Model/Data \ Compromise:} & J_n(v) = \frac{1}{2}|v - \widehat{v}_{n+1}|_{\widehat{C}}^2 + \frac{1}{2}|y_{n+1} - Hv|_{\Gamma}^2\\ {\sf Optimize:} & v_{n+1} = {\rm argmin}_v \ J_n(v). \end{array}$$

- \widehat{C} is a fixed model covariance (not updated sequentially).
- $d = \mathcal{O}(10^9)$; $\mathcal{O}(d^2)$ entries of \widehat{C} prohibitive in general.
- \widehat{C} based on climatology + simple, computable, structure.

Ensemble Kalman Filter (Oceanography)

State Space Model

Dynamics Model: $v_{n+1} = \Psi(v_n) + \xi_n$, $n \in \mathbb{Z}^+$ Data Model: $y_{n+1} = Hv_{n+1} + \eta_{n+1}$, $n \in \mathbb{Z}^+$ Probabilistic Structure: $v_0 \sim N(m_0, C_0)$, $\xi_n \sim N(0, \Sigma)$, $\eta_n \sim N(0, \Gamma)$ Probabilistic Structure: $v_0 \perp \{\xi_n\} \perp \{\eta_n\}$ independent



- Geir Evensen.
- ▶ See [11].
- Motivated by extended Kalman filter; see [17, 14].
- Original paper in oceanography.
- Now used in weather forecasting centres worldwide.

•
$$\{v_n^{(j)}\}_{j=1}^J \mapsto \{v_{n+1}^{(j)}\}_{j=1}^J$$

Ensemble Kalman Filter

Sequential Optimization Perspective

$$\begin{array}{ll} {\sf Predict:} & \widehat{v}_{n+1}^{(j)} = \Psi(v_n^{(j)}) + \xi_n^{(j)}, & n \in \mathbb{Z}^+ \\ {\sf Model/Data \ Compromise:} & J_n^{(j)}(v) = \frac{1}{2} |v - \widehat{v}_{n+1}^{(j)}|_{\widehat{\mathcal{C}}_{n+1}}^2 + \frac{1}{2} |y_{n+1} - Hv|_{\Gamma}^2 \\ {\sf Optimize:} & v_{n+1}^{(j)} = \operatorname{argmin}_v \ J_n^{(j)}(v). \end{array}$$

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- ▶ $j \in \{1, ..., J\}$, J number of ensemble members.
- \widehat{C}_{n+1} is empirical covariance of the $\{\widehat{v}_{n+1}^{(k)}\}$.
- Updating \widehat{C}_n requires only $\mathcal{O}(Jd)$ storage.

Ensemble Kalman Filter (Mathematical Structure)

State Space Model

Dynamics Model: $v_{n+1} = \Psi(v_n) + \xi_n$, $n \in \mathbb{Z}^+$ Data Model: $y_{n+1} = Hv_{n+1} + \eta_{n+1}$, $n \in \mathbb{Z}^+$ Probabilistic Structure: $v_0 \sim N(m_0, C_0)$, $\xi_n \sim N(0, \Sigma)$, $\eta_n \sim N(0, \Gamma)$ Probabilistic Structure: $v_0 \perp \{\xi_n\} \perp \{\eta_n\}$ independent



- Sebastian Reich (Potsdam)
- Continuous Time Limits:
- ▶ [4, 5, 27, 6].
- Optimal Transport Connections:
- ▶ [27, 29, 28].
- SFB 1294 (Potsdam)

Weather Forecasting

3DVAR Overcomes Butterfly Effect KJH Law and S [24].





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3DVAR Overcomes Butterfly Effect KJH Law and S [24].





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Impact of EnKF over 3DVAR

Courtesy of Roland Potthast (Head of Data Assimilation, DWD)



Ensemble Kalman Inversion

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- Chen & Oliver [8]
- Emerick and Reynolds [9]
- Iglesias, Law and S [16]
- Ernst, Sprungk and Starkloff [10]

Inverse Problem

Problem Statement

Find \boldsymbol{u} from \boldsymbol{y} where $\mathsf{G}:\mathcal{U}\mapsto\mathcal{Y},\,\eta$ is noise and

 $y = \mathsf{G}(\mathbf{u}) + \eta.$

Main Approaches

$$\begin{array}{ll} \textit{Optimization} \quad \Phi(u) = \frac{1}{2}|y - G(u)|_{\Gamma}^2 + \frac{1}{2}|u|_{\Sigma}^2;\\ \textit{Probability} \quad \mathbb{P}(u|y) \propto \exp(-\Phi(u)). \end{array}$$

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Inverse Problem

Dynamical Formulation



- Evensen moved to Statoil.
- Dean Oliver & Al Reynolds; see [8, 9].
- $y_{n+1} = y$, $\eta_{n+1} \sim N(0, M\Gamma)$.
- Methodology now widely used in oil industry.
- Methodology now widely used by hydrologists.

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Discrete Time Kalman Inversion Algorithm

Covariances

$$\begin{split} C_n^{ww} &= \frac{1}{J} \sum_{j=1}^J \left(\mathsf{G}(u_n^{(j)}) - \overline{w}_n \right) \otimes \left(\mathsf{G}(u_n^{(j)}) - \overline{w}_n \right), \quad \overline{w}_n = \frac{1}{J} \sum_{j=1}^J \mathsf{G}(u_n^{(j)}), \\ C_n^{uw} &= \frac{1}{J} \sum_{j=1}^J \left(u_n^{(j)} - \overline{u}_n \right) \otimes \left(\mathsf{G}(u_n^{(j)}) - \overline{w}_n \right), \quad \overline{u}_n = \frac{1}{J} \sum_{j=1}^J u_n^{(j)}. \end{split}$$

Iteration $n \mapsto n+1$

$$u_{n+1}^{(j)} = u_n^{(j)} + C_n^{uw} (C_n^{ww} + M\Gamma)^{-1} (y - G(u_n^{(j)}))$$

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Discrete Time Kalman Inversion Algorithm

Covariances

$$\begin{split} C_n^{ww} &= \frac{1}{J} \sum_{j=1}^J \left(\mathsf{G}(u_n^{(j)}) - \overline{w}_n \right) \otimes \left(\mathsf{G}(u_n^{(j)}) - \overline{w}_n \right), \quad \overline{w}_n = \frac{1}{J} \sum_{j=1}^J \mathsf{G}(u_n^{(j)}), \\ C_n^{uw} &= \frac{1}{J} \sum_{j=1}^J \left(u_n^{(j)} - \overline{u}_n \right) \otimes \left(\mathsf{G}(u_n^{(j)}) - \overline{w}_n \right), \quad \overline{u}_n = \frac{1}{J} \sum_{j=1}^J u_n^{(j)}. \end{split}$$

Iteration $n \mapsto n+1$

$$u_{n+1}^{(j)} = u_n^{(j)} + C_n^{uw} (C_n^{ww} + M\Gamma)^{-1} (y - \mathsf{G}(u_n^{(j)}))$$

Continuous Time Limit

$$u_n^{(j)} \approx u^{(j)}(t)|_{t=n/M} : \dot{\boldsymbol{u}}^{(j)} = -\frac{1}{J} \sum_{k=1}^{J} \left\langle \mathsf{G}(\boldsymbol{u}^{(k)}) - \bar{\mathsf{G}}, \mathsf{G}(\boldsymbol{u}^{(j)}) - \boldsymbol{y} \right\rangle_{\Gamma} \left(\boldsymbol{u}^{(k)} - \bar{\boldsymbol{u}} \right)$$

Gradient Flow In Parameter Space

- Ensemble Filtering Continuous Time: Bergemann & Reich (2010a, 2010b, 2012) [4, 5, 6]
- Ensemble Filtering Continuous Time: Reich (2011) [27]
- Connection to Foais/Prodi: Titi and coworkers [15, 2]
- 3DVAR Filtering Continuous Time: Blömker, Law, S & Zygalakis (2013) [7]
- Ensemble Filtering Continuous Time: Kelly, Law & S (2015) [20]
- Ensemble Inversion Continuous Time: Schillings & S (2017) [30]
- Text: Reich & Cotter (2015) [28]
- Text: Law, S & Zygalakis (2015) [23]
- Ensemble Filtering Continuous Time: Lange & Stannat [22]
- Ensemble Square Root Filtering Continuous Time: Lange & Stannat [21]

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Gradient Flow In Space Of Probability Measures

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- Jordan, Kinderlehrer & Otto 1998 [18]
- Otto 2001 [26]
- Benamou & Brenier 2000 [3]
- Ambrosio, Gigli & Savare 2008 [1]
- Villani 2008 [31]
- Reich & Cotter 2013 [29]
- Garbuno-Inigo, Hoffmann, Li & Stuart 2020 [12]
- Garbuno-Inigo, Nüsken & Reich [13]

Ensemble Kalman Inversion (EKI) – Linear G see [27], [30]

EKI Is Self-Preconditioned Gradient Descent

$$\begin{split} \dot{\boldsymbol{\mu}}^{(j)} &= -C(\boldsymbol{u}) \nabla \Phi_0(\boldsymbol{u}^{(j)}), \quad \Phi_0(\boldsymbol{u}) = rac{1}{2} \big| \boldsymbol{y} - G(\boldsymbol{u}) \big|_{\Gamma}^2, \ &ar{\boldsymbol{u}} = rac{1}{J} \sum_{k=1}^J \boldsymbol{u}^{(k)}, \quad C(\boldsymbol{u}) = rac{1}{J} \sum_{k=1}^J \Big(\boldsymbol{u}^{(k)} - ar{\boldsymbol{u}} \Big) \otimes \Big(\boldsymbol{u}^{(k)} - ar{\boldsymbol{u}} \Big). \end{split}$$

Theorem [30]

EKI minimizes Φ_0 over a finite dimensional subspace determined by the initial conditions $\{ \boldsymbol{u}^{(j)}(0) \}_{j=1}^{j}$. The rate of convergence is $\mathcal{O}(1/t)$.

Ensemble Kalman Sampling – Linear G see [12]

EKS Is Self-Preconditioned Langevin Equation

$$\begin{split} \dot{u}^{(j)} &= -C(u) \nabla \Phi(u^{(j)}) + \sqrt{2C(u)} \dot{W}^{(j)}, \quad \Phi(u) = \frac{1}{2} |y - G(u)|_{\Gamma}^{2} + \frac{1}{2} |u|_{\Sigma}^{2}, \\ \bar{u} &= \frac{1}{J} \sum_{k=1}^{J} u^{(k)}, \quad C(u) = \frac{1}{J} \sum_{k=1}^{J} \left(u^{(k)} - \bar{u} \right) \otimes \left(u^{(k)} - \bar{u} \right). \end{split}$$

Mean Field Limit: Nonlinear Nonlocal Fokker-Planck Eq.

$$\begin{split} \dot{\boldsymbol{u}} &= -\mathcal{C}(\rho)\nabla\Phi(\boldsymbol{u}) + \sqrt{2C(\rho)}\dot{\boldsymbol{W}},\\ \mathcal{C}(\rho) &= \int \left(\boldsymbol{u} - \bar{\boldsymbol{u}}\right)\otimes\left(\boldsymbol{u} - \bar{\boldsymbol{u}}\right)\rho(\boldsymbol{u},t)d\boldsymbol{u}, \quad \bar{\boldsymbol{u}} = \int \boldsymbol{u}\rho(\boldsymbol{u},t)d\boldsymbol{u},\\ \partial_t\rho &= \nabla\cdot\left(\rho\,\mathcal{C}(\rho)\,\nabla\Phi\right) + \mathcal{C}(\rho):D^2\rho, \quad \rho(0) = \rho_0. \end{split}$$

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Nonlinear Nonlocal Fokker-Planck Equation See [18], [12]

Theorem [12]

The nonlinear Fokker-Planck equation may be written as

$$\partial_t \rho = \nabla \cdot \left(\rho \, \mathcal{C}(\rho) \nabla \frac{\delta \mathcal{E}}{\delta \rho} \right) \ , \ \mathcal{E}(\rho) = \int \left(\Phi + \ln \rho \right) \rho \, \mathrm{d} u.$$

- Gradient flow in P₊ (probability measures) w.r.t. the (next slide) Kalman-Wasserstein metric;
- in linear setting convergence to equilibrium ∝ exp(-Φ) occurs at exponential rate exp(-t), indepedently of the linear inverse problem being solved.

Nonlinear Nonlocal Fokker-Planck Equation (Metric) [29].[12]

Kalman-Wasserstein Metric Tensor (Otto [18, 26]) Define $g_{\rho, \mathcal{C}}$: $T_{\rho}\mathcal{P}_{+} \times T_{\rho}\mathcal{P}_{+} \to \mathbb{R}$ by $g_{\rho, \mathcal{C}}(\sigma_{1}, \sigma_{2}) := \int_{\Omega} \langle \nabla \psi_{1}, \mathcal{C}(\rho) \nabla \psi_{2} \rangle \rho \, \mathrm{d}x,$ where $\sigma_{i} = -\nabla \cdot (\rho \mathcal{C}(\rho) \nabla \psi_{i}) \in T_{\rho}\mathcal{P}_{+}$ for i = 1, 2. Kalman-Wasserstein Metric (Benamou-Brenier [3])

For ρ^0 , $\rho^1 \in \mathcal{P}_+$, $\mathcal{W}_{\mathcal{C}} \colon \mathcal{P}_+ \times \mathcal{P}_+ \to \mathbb{R}$

$$\mathcal{W}_{\mathcal{C}}(\rho^{0},\rho^{1})^{2} := \inf_{(\rho_{t},\psi_{t})} \int_{0}^{1} \int_{\Omega} \langle \nabla \psi_{t} , \mathcal{C}(\rho_{t}) \nabla \psi_{t} \rangle \rho_{t} \, \mathrm{d}x$$

subject to $\partial_{t} \rho_{t} + \nabla \cdot (\rho_{t} \mathcal{C}(\rho_{t}) \nabla \psi_{t}) = 0, \ \rho_{0} = \rho^{0}, \ \rho_{1} = \rho^{1},$

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Closing

Conclusions: Kalman Methodologies

- Introduced in 1960 by Rudolh Kalman.
- Basic algorithm generalized in many directions.
- Applications in numerous fields:
 - Navigation;
 - Weather forecasting;
 - Oceanography;
 - Hydrology, Subsurface flow;
 - Medical imaging, Machine learning · · · .
- Developing as a general methodology for state estimation.
- Developing as a general methodology for inverse problems:
 - Gradient flow structure: parameter space;
 - Gradient flow structure: probability space.
- Connections to Wasserstein gradient flows, optimal transport.

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Many open mathematical questions.

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Ensemble Kalman Inversion (EKI)

Continuous Time Formulation $\dot{\boldsymbol{u}}^{(j)} = -\frac{1}{J} \sum_{k=1}^{J} \left\langle \mathsf{G}(\boldsymbol{u}^{(k)}) - \bar{\mathsf{G}}, \mathsf{G}(\boldsymbol{u}^{(j)}) - \boldsymbol{y} \right\rangle_{\Gamma} \left(\boldsymbol{u}^{(k)} - \bar{\boldsymbol{u}} \right)$ $\bar{\boldsymbol{u}} = \frac{1}{J} \sum_{k=1}^{J} \boldsymbol{u}^{(k)}, \quad \bar{\mathsf{G}} = \frac{1}{J} \sum_{k=1}^{J} \boldsymbol{G}(\boldsymbol{u}^{(k)}),$ $\boldsymbol{C}(\boldsymbol{u}) = \frac{1}{J} \sum_{k=1}^{J} \left(\boldsymbol{u}^{(k)} - \bar{\boldsymbol{u}} \right) \otimes \left(\boldsymbol{u}^{(k)} - \bar{\boldsymbol{u}} \right).$

Linear Approximation

$$(\mathsf{G}(\boldsymbol{u}^{(k)})-\bar{\mathsf{G}})\approx D\mathsf{G}(\boldsymbol{u}^{(j)})(\boldsymbol{u}^{(k)}-\bar{\boldsymbol{u}}).$$

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Ensemble Kalman Sampling (EKS)

Continuous Time Formulation: Put EKI in a heat bath

$$\begin{split} \dot{\boldsymbol{u}}^{(j)} &= -\frac{1}{J} \sum_{k=1}^{J} \left\langle \mathsf{G}(\boldsymbol{u}^{(k)}) - \bar{\mathsf{G}}, \mathsf{G}(\boldsymbol{u}^{(j)}) - \boldsymbol{y} \right\rangle_{\Gamma} \left(\boldsymbol{u}^{(k)} - \bar{\boldsymbol{u}} \right) \\ &- C(\boldsymbol{u}) \boldsymbol{\Sigma}^{-1} \boldsymbol{u}^{(j)} + \sqrt{2C(\boldsymbol{u})} \dot{\boldsymbol{W}}^{(j)}, \\ C(\boldsymbol{u}) &= \frac{1}{J} \sum_{k=1}^{J} \left(\boldsymbol{u}^{(k)} - \bar{\boldsymbol{u}} \right) \otimes \left(\boldsymbol{u}^{(k)} - \bar{\boldsymbol{u}} \right). \end{split}$$

Gradient Structure Of NNLFP in MFL

$$\begin{aligned} \frac{\mathrm{d}}{\mathrm{d}t} \mathcal{E}(\rho) &= -\int \rho \left| \mathcal{C}(\rho)^{\frac{1}{2}} \nabla (\Phi + \ln \rho) \right|^2 \mathrm{d}\boldsymbol{u} \\ &= -\boldsymbol{g}_{\rho,\mathcal{C}}(\partial_t \rho, \partial_t \rho). \end{aligned}$$

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